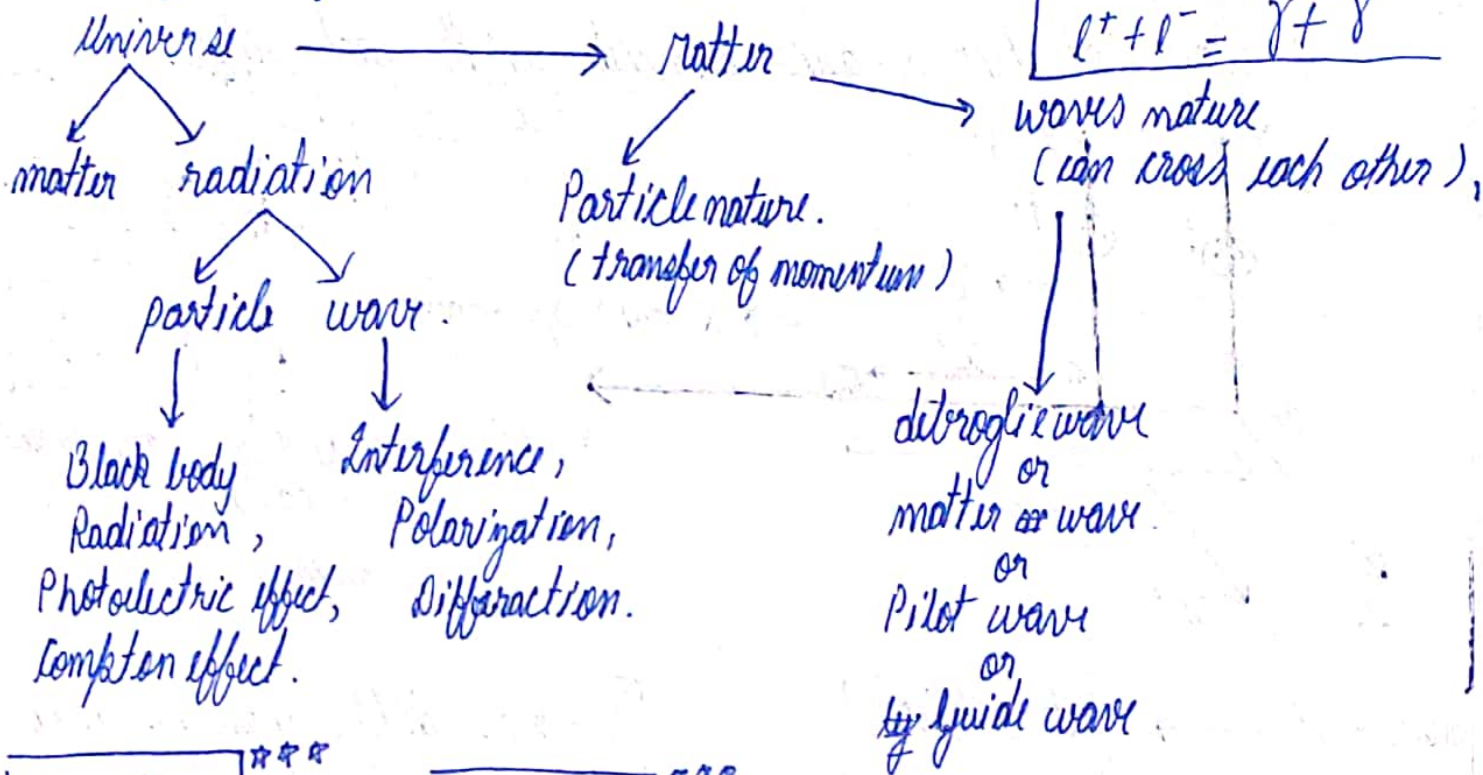


Quantum mechanics:-

(Bharat deep Singh, 26, CSE-A1)

De-Broglie's Hypothesis (Matter waves)



$$\gamma = l^- + l^+$$

$$l^+ + l^- = \gamma + \gamma$$

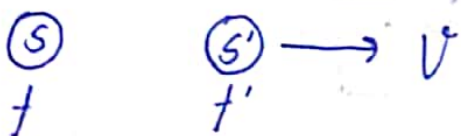
$$\lambda = \frac{h}{m v}$$

$$E = m c^2$$

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}}$$

$$K \cdot E = (m - m_0) c^2$$

$$U = m_0 c^2$$



$$t' = t - \frac{v x}{c^2 \sqrt{1 - \frac{v^2}{c^2}}}$$

angular freq. $k = \frac{2\pi}{\lambda}$

$$\Psi_p = A \sin(\omega t - kx)$$

Amplitude $\omega = 2\pi \mu$

propagation const.

Davison
and
Germer
Experiment:-

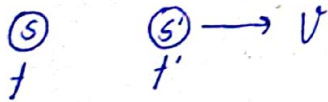
$$\lambda = \frac{h}{mv}$$

$$E = mc^2$$

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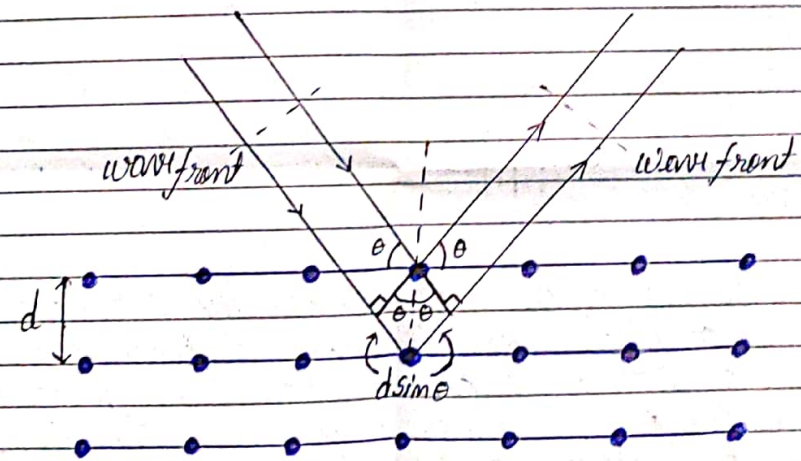
$$\Psi_p = A \sin(\omega t - kx)$$

Amplitude $= 2\pi A$

propagation const.

* Davission and Germer Experiment :-

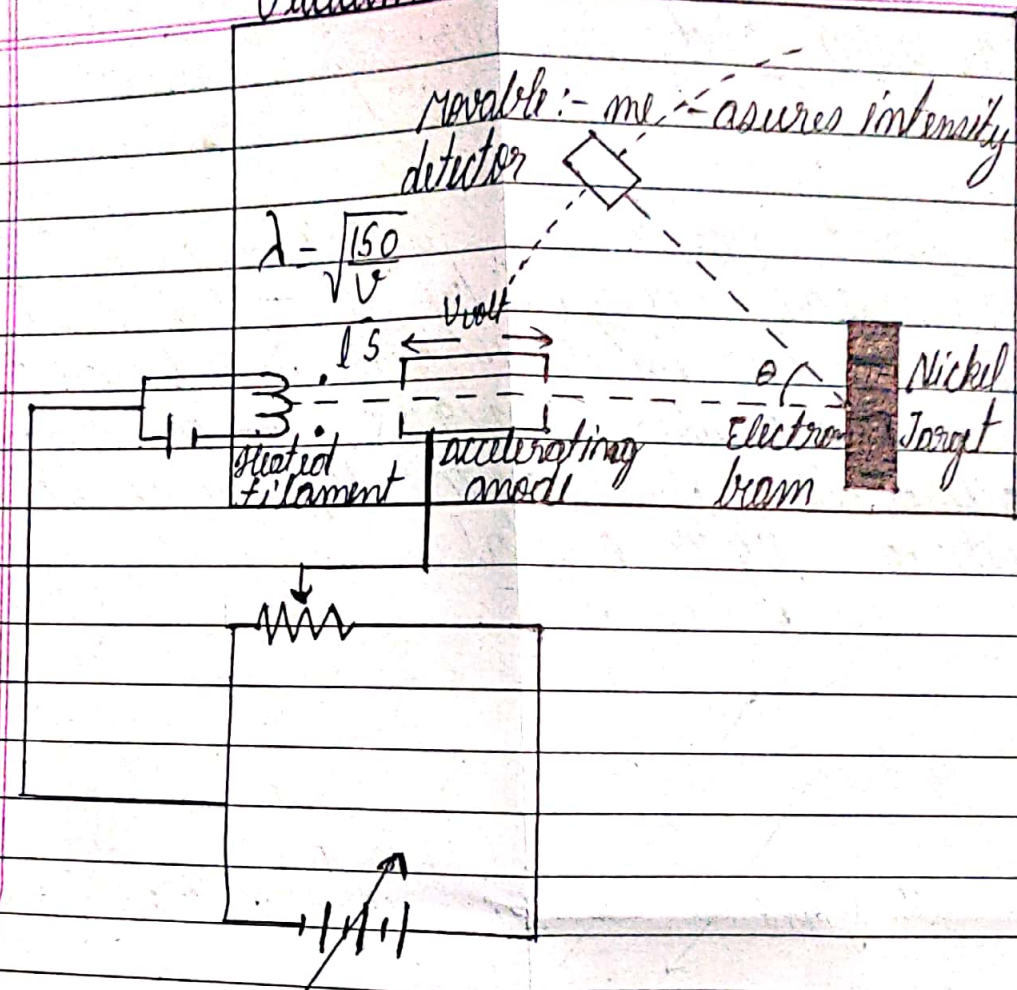
The wave nature of electrons was first experimentally verified by C.T. Davission and J.H. Germer in 1927 and independently by G.P. Thomson, in 1928, who observed diffraction effects with beams of electrons scattered by crystal.



$$\Delta x = 2d \sin \theta = m \lambda \quad (\text{Constructive interference})$$

\Rightarrow high intensity e^- beam

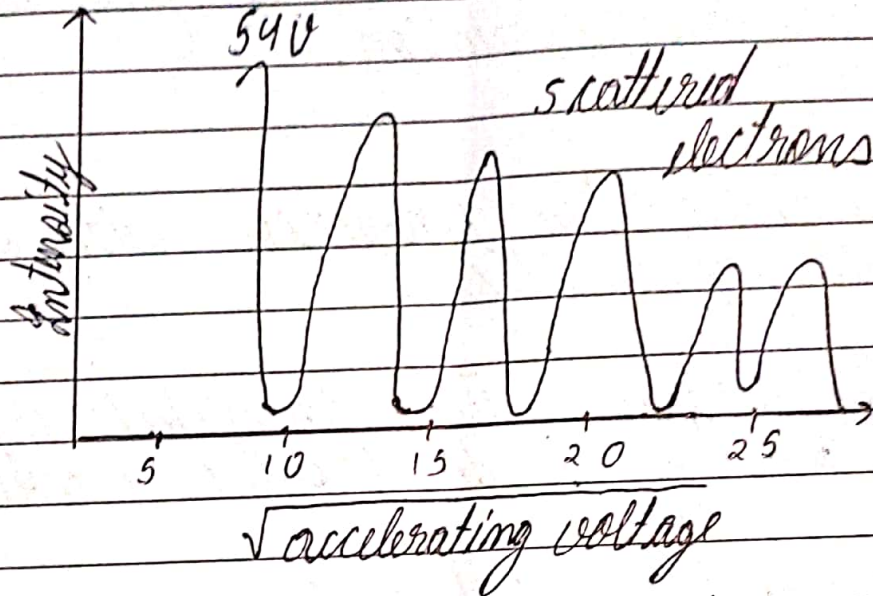
Vacuum chamber



By moving the detector on the circular scale at diff. positions, the intensity of the scattered electron beam is measured for diff. values of angle θ w.r. to the incident and scattered e-beams.

The variation of Intensity (I) of the scattered electrons with the angle of scattering is obtained from diff. accelerating voltages.

* graph:-



* The experiment was performed by varying the accelerating voltage from 44V to 68V. It was noticed that a strong peak appeared in the intensity (I) of the scattered electron for an accelerated voltage of 54V at a scattering angle $\theta = 50^\circ$.

Result:-

From the electron diffraction measurements the wave-length of matter waves was found to be 0.165nm

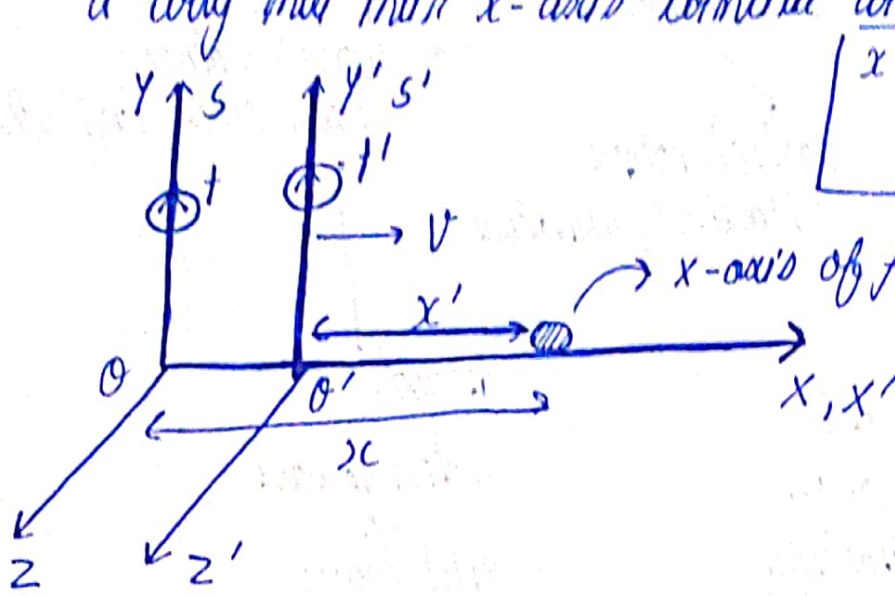
The de-Broglie wavelength λ associated with electrons for $V = 54V$ is given by

$$\lambda = \frac{h}{p} = \frac{1.227 \text{ nm}}{\sqrt{V}}$$

$$\lambda = \frac{1.227}{\sqrt{54}} = 0.167 \text{ nm}$$

$$\Psi_s = A \sin(\omega t) \cos(kx)$$

Prove: Suppose there are two frames of reference, S and S' in such a way that their x-axes coincide with each other.



$$\begin{cases} x' = \frac{x - vt}{\sqrt{1 - \frac{v^2}{c^2}}} \\ y' = y, z' = z \\ t' = \frac{t - \frac{v}{c^2}x}{\sqrt{1 - \frac{v^2}{c^2}}} \end{cases}$$

These 4 eqn. are collectively called Lorentz Transformation.

$m_0 \leftarrow S' \rightarrow$ rest (object)
 $m \leftarrow S \rightarrow$ motion (object) $\rightarrow v$
 \downarrow
 mass of objects acc. to frame of reference.

$\rightarrow \therefore$ diff. mass in diff. frame of reference.

$x' \rightarrow$ const.
 $x \rightarrow$ variable.

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} \quad \text{--- ①}$$

We have assumed that the object possesses the wave nature
 \therefore wave is observable in frame 'S' and 'S'

$$\left| \Psi'_s = A' \sin(\omega' t') \cos(k' x') \right| \quad \text{--- ②}$$

\downarrow some wave seen from 'S'
 \downarrow stationary wave.
 must appear to be progressive wave

$$\left| \Psi'_s = A \sin(\omega' t') \right| \quad \text{--- ③}$$

($\because A = A' \cos(k' z')$)

$$t' = t - \frac{v}{c^2} x \quad \rightarrow \textcircled{4} \quad \psi' = \psi$$

$$\Rightarrow \psi = A \sin \left[\omega' \left(t - \frac{v}{c^2} x \right) \right]$$

in the frame 'S'

$$\psi = A \sin \left[\omega \left(t - \frac{v}{c^2} x \right) \right] \quad \text{---} \textcircled{5}$$

$$\therefore \omega = \frac{\omega'}{\sqrt{1 - \frac{v^2}{c^2}}}$$

$$\psi = A \sin \left[\omega t - \frac{\omega v}{c^2} x \right]$$

$$\psi_p = A \sin(\omega t - kx) \quad \text{---} \textcircled{6} \quad \left[\because k = \frac{\omega v}{c^2} \right]$$

$v_p =$ phase velocity or wave velocity. \rightarrow $\frac{\text{coeff of 't'}}{\text{coeff of 'x'}}$

$$= v \lambda$$

$$= \frac{2\pi v \lambda}{2\pi} = \frac{2\pi v}{\frac{2\pi}{\lambda}} = \frac{\omega}{k}$$

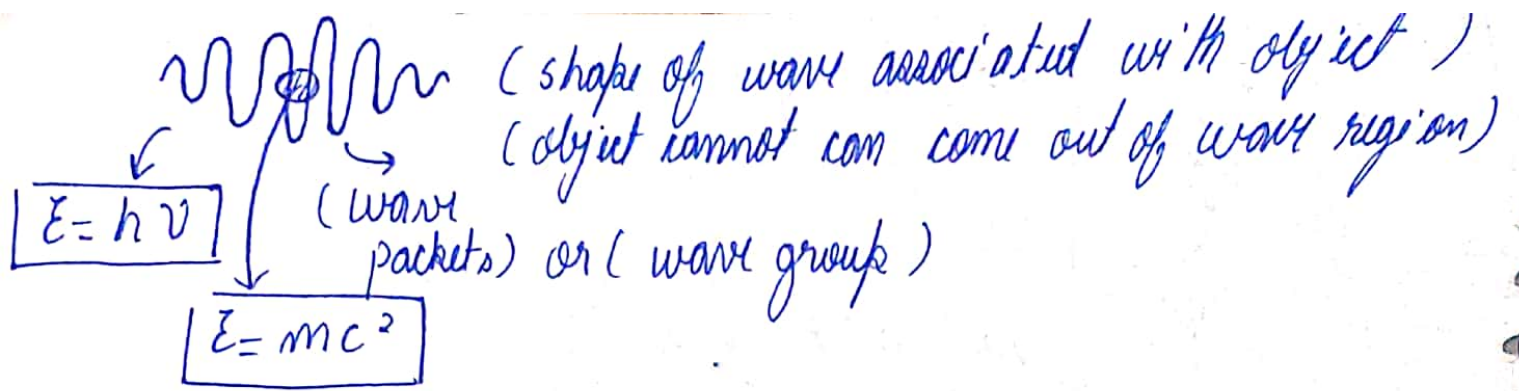
$$\therefore v_p = \frac{\omega}{k} \quad \text{---} \textcircled{7}$$

$$v_p = \frac{\omega c^2}{\omega v} \quad \rightarrow \quad v_p = \frac{c^2}{v} \quad \text{---} \textcircled{8}$$

$$v_p = c \left(\frac{c}{v} \right) \quad v_p > c \text{ (always)}$$

always greater than '1'

\therefore This means de Broglie waves travel faster than light in vacuum



Since the wave packet belongs to the particle only. \therefore Energy of the wave packet and energy of the object which is confined inside the wave packet must be same

$$E = h\nu \longrightarrow E = \frac{h\nu_p}{\lambda} \longrightarrow E = \frac{h}{\lambda} \frac{c^2}{v} \quad \text{--- (9)}$$

[\because (8)]

$$\therefore mc^2 = \frac{h}{\lambda} \frac{c^2}{v}$$

$$\boxed{\lambda = \frac{h}{mv}} \longrightarrow \boxed{\lambda = \frac{h}{p}}$$

mass in motion. (10)

We obtained this result by using the laws of "theory of relativity" so, \therefore eqn. (10) is valid at low speeds as well as high speeds. \therefore eqn. (10) is a relativistic expression.

Let us take an electron :-

$$m_0 = 9.1 \times 10^{-31} \text{ kg.}$$

$$v = 1.6 \times 10^8 \text{ m/s.}$$

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}}$$

$$m = \frac{m_0}{\sqrt{1 - \frac{1}{4}}}$$

$$m = \frac{2m_0}{\sqrt{3}} = \frac{2m_0}{1.732}$$

$$m = \frac{2}{1.732} \times 9.1 \times 10^{-31} = \frac{18.2}{1.732} \times 10^{-31} \text{ kg.}$$

$$\lambda = \frac{h}{m v} = \frac{6.63 \times 10^{-34} \text{ J s}}{\frac{18.2}{1.732} \times 10^{-31} \times 1.5 \times 10^6} \approx 0.4 \times 10^{-11} \text{ m}$$

* Wave function $\rightarrow (\Psi)$ (changes its value w.r.t. to space & time)

A wave fn. can be defined as a probability amplitude which is given in quantum mechanics, to define stats of a particle and see how the particle behave.

eg:- Light waves \rightarrow Electric, Magnetic fields which repeats its value again and again w.r.t. space and time

$\Psi \rightarrow$ ~~stat~~ Electric or magnetic field

eg:- sound waves \rightarrow (Pressure or Density of medium)

eg:- Matter waves \rightarrow (cannot give physical significance to wave fn Ψ)

whereas $|\Psi|^2 \rightarrow$ measurable quantity.

not a measurable quantity but a calculative quantity

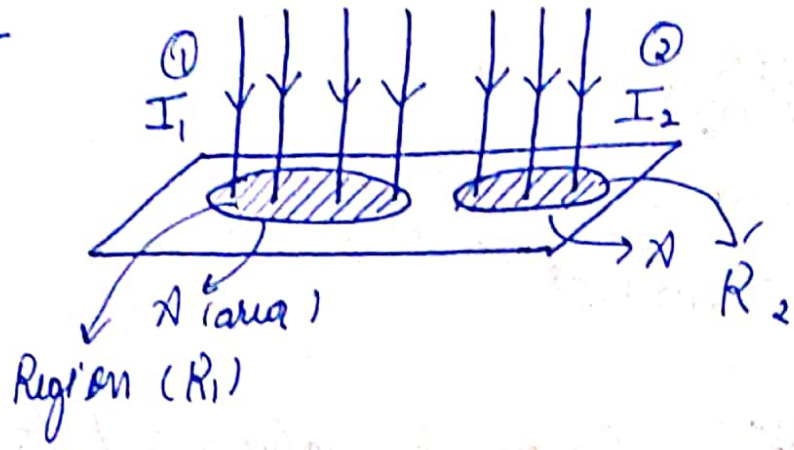
acc. to Max Born in a small region around any pt. $\therefore |\Psi|^2$ will represent probability of finding the particle whose matter waves are being discussed in a unit region drawn around that pt.



$$|\Psi|^2 \propto \frac{dP}{dV}$$

(dP = small Probability of finding a p article in this small region of volume dV)

eg:-



finding photon in R_1 and R_2

due to $I_1 > I_2$ the probability of finding a photon in $R_1 > R_2$

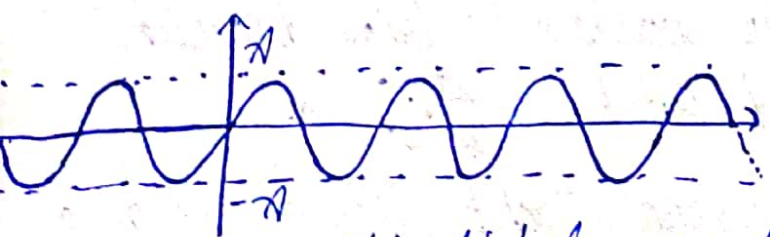
\therefore finding photon probability \propto intensity $\propto A^2$
 (Amplitude)
 $A^2 = |\Psi|^2 \propto \frac{dP}{dV}$ or $\frac{dP}{dA}$ or $\frac{dP}{dx}$
 \downarrow \downarrow \downarrow
 3-D 2-D 1-D

Nature of de-Broglie waves

$$\lambda = \frac{h}{mv}$$

$$\Psi = A \sin(\omega t - kx) \quad (A = \text{const.}) \quad (\text{Plane progressive wave.})$$

range $[-A, A]$
 Domain all real numbers
 continuous fn.
 +ve x-axis.



$$A^2 \rightarrow |\Psi|^2 \propto \frac{dP}{dx}$$

1-D

This kind of wave fn cannot be attached to a particle because the particle cannot be present everywhere in the universe at an instant.

2nd reason:- $v_p = \frac{c^2}{v}$ ($v = \text{speed of object}$)

$v_p = \text{speed of matter wave associated with object}$

$\frac{v_p}{c} = \frac{c}{v}$
 not bounded. bounded

$v_p > c$
 $v < c$

This shows that de-Broglie wave will detach from object.

Suppose Two plane progressive wave.

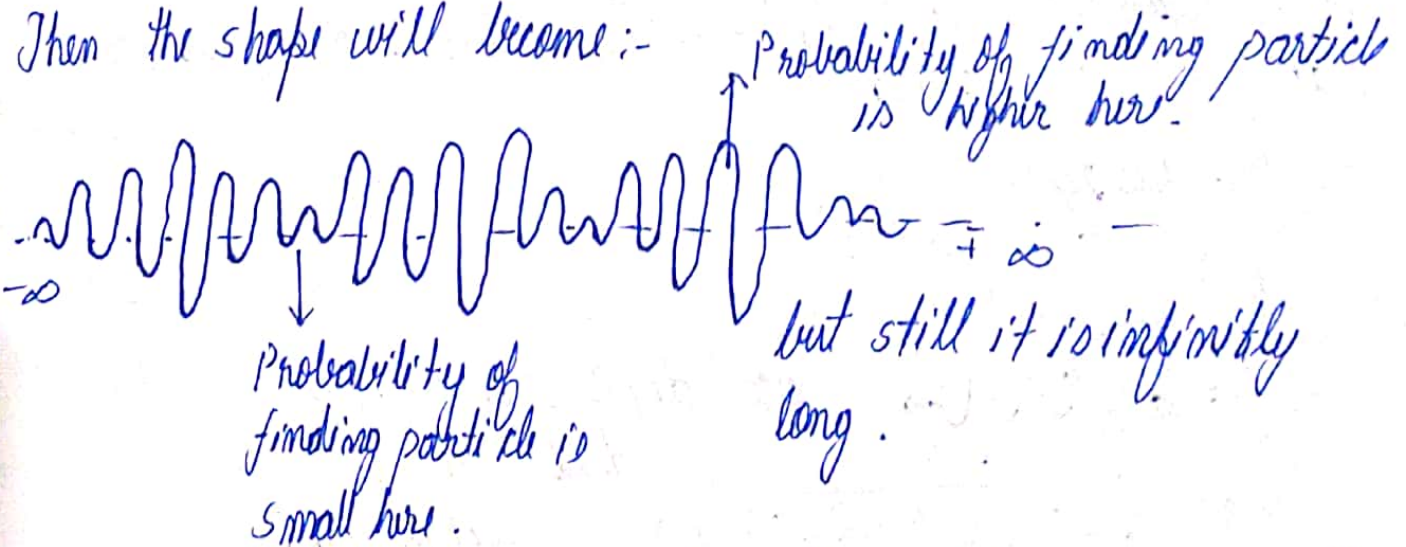
$\Psi_1 = A_1 \sin(\omega_1 t - k_1 x)$ — ①

$\Psi_2 = A_2 \sin(\omega_2 t - k_2 x)$ — ②

Superimposing ① & ②

$\Psi_1 + \Psi_2 = ?$ (Somewhere the interference will be constructive and somewhere destructive)

Then the shape will become:-



first the particle was present in the whole universe with equal probability but after superimposing the waves the particle

are still present in the wave universal but with unequal probabilities

un-desired

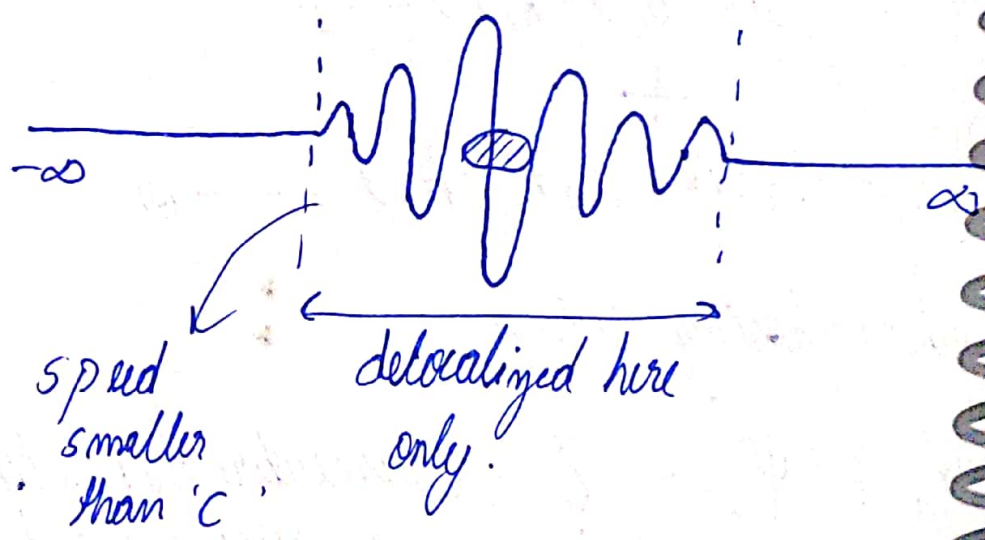
desired

* If we add more ^{and} more progressive waves then the length of the wave will start shrinking

i.e. $\psi_1 + \psi_2 + \psi_3 + \psi_4 + \dots = \psi$ (wave packet or wave group)

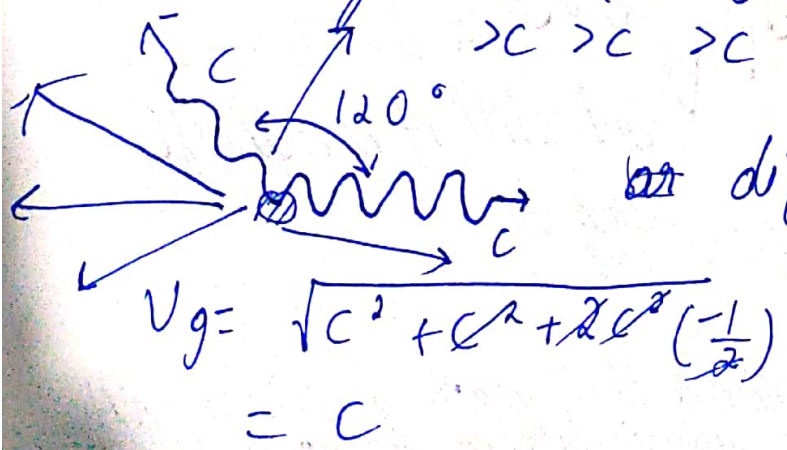
∴ a wave packet or wave group must be attached with an object instead of a de-broglie wave.

∴ we will get



It has been formed by superimposing of waves with speeds $U_1, U_2, U_3, \dots, U_n$ all greater than 'c' → velocities.

∴ but always $U_1 + U_2 + U_3 + \dots + U_n = U_g$
 $> c > c > c \dots > c < c$



or diff. direction diff waves vector addition

Formation of wave packet :->

Power & group
N/A this
part the re.
and linear
d'Alambert.

$$\psi_1 = A \sin(\omega t - kx) \quad \text{--- ①}$$

$$\psi_2 = A \sin((\omega + d\omega)t - (k + dk)x) \quad \text{--- ②}$$

$$\psi = \psi_1 + \psi_2$$

$$= A \left\{ \underbrace{\sin[(\omega + d\omega)t - (k + dk)x]}_C + \underbrace{\sin(\omega t - kx)}_D \right\}$$

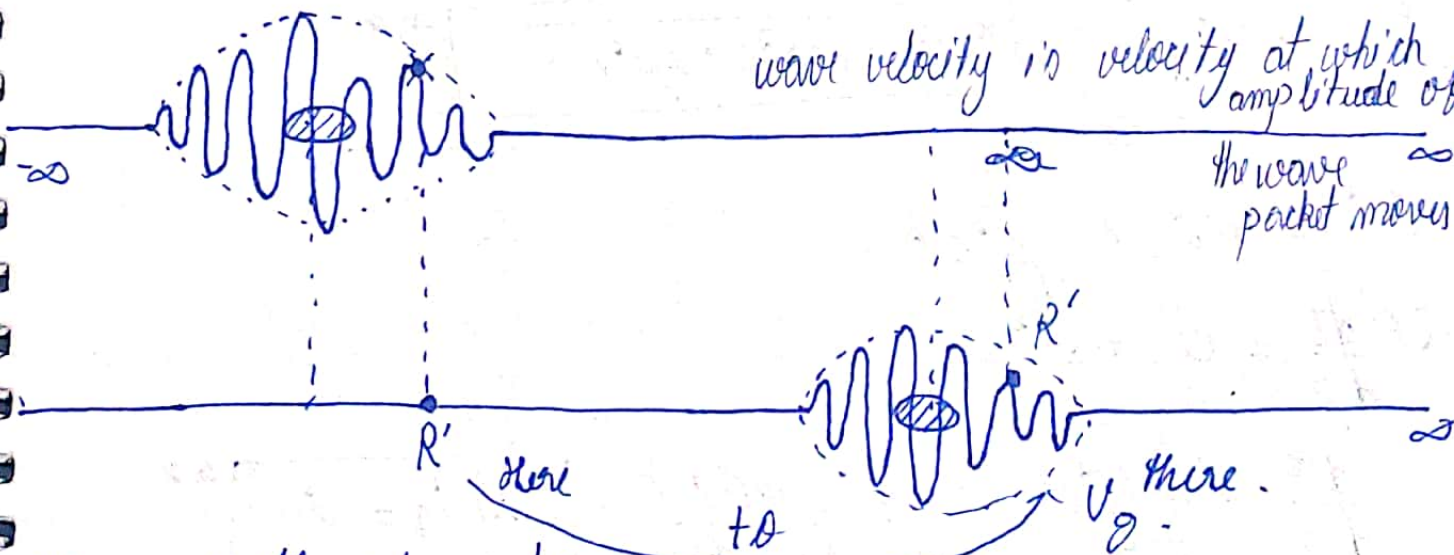
$$= 2A \sin\left[(\omega + \frac{d\omega}{2})t - (k + \frac{dk}{2})x\right] \cos\left[\frac{d\omega}{2}t - \frac{dk}{2}x\right] \quad \text{--- ③}$$

Put $2A \cos\left(\frac{d\omega}{2}t - \frac{dk}{2}x\right) = R \quad \text{--- ④}$

$d\omega$, and dk are small.

$$\psi \approx R \sin(\omega t - kx) \quad \text{--- ⑤} \quad \text{[Amplitude is variable [} R = f(t, x) \text{]}$$

wave velocity is velocity at which amplitude of the wave packet moves.



$$v_{gR} = \frac{\text{coeff. of } t}{\text{coeff. of } x} = \frac{\frac{d\omega}{2}}{\frac{dk}{2}} = \frac{d\omega}{dk}$$

$$\therefore \boxed{v_g = \frac{d\omega}{dk}} \quad \text{--- (6)}$$

$$v_{gr} = \frac{\omega + d\omega}{k + dk} - \frac{\omega}{k} = v_g = \frac{d\omega}{dk}$$

we know that: $E = h\nu = \frac{h \times 2\pi\nu}{2\pi} = \hbar\omega$

$$\therefore \boxed{E = \hbar\omega} \quad \text{--- (7)} \quad \hbar = 1.054 \times 10^{-34} \text{ J s}$$

↓
Planck's reduced const.

$$p = \frac{h}{\lambda} = \frac{h}{2\pi} \times \frac{2\pi}{\lambda} = \hbar k$$

$$\Rightarrow \boxed{p = \hbar k} \quad \text{--- (8)}$$

7, 8 are called Einstein's de Broglie's eqs.

$$v_g = \frac{d\omega}{dk} = \frac{\hbar}{\hbar} \frac{d\omega}{dk} = \frac{d(\hbar\omega)}{d(\hbar k)}$$

$$\boxed{v_g = \frac{dE}{dp}} \quad \text{--- (9)}$$

$$\boxed{E = mc^2} \quad \text{--- (10)}$$

$$\boxed{E = \sqrt{m_0^2 c^4 + p^2 c^2}} \quad \text{--- (11)}$$

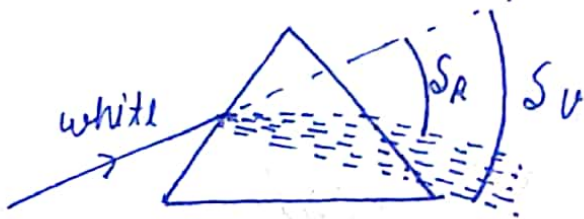
$$E^2 = m_0^2 c^4 + p^2 c^2$$

$$2E \frac{dE}{dp} = 0 + 2c^2 p$$

$$\frac{dE}{dp} = \frac{2c^2 p}{2E} = \frac{c^2 (mv)}{mc^2}$$

$$\therefore \boxed{v_g = v} \quad \text{--- (12)}$$

Dispersion relation for de-broglie waves:-



$$\mu = c_1 + \frac{c_2}{\lambda^2} + \frac{c_3}{\lambda^3} \dots \dots \dots \infty$$

$$\left. \begin{array}{l} \text{speed of} \\ \text{light in} \\ \text{medium} \end{array} \right\} = \frac{c}{\mu}$$

$$\mu \approx c_1 + \frac{c_2}{\lambda^2}, \text{ as } \lambda \uparrow \Rightarrow \mu \downarrow \Rightarrow \text{Speed} \uparrow$$

$$S = (\mu - 1) A$$

We know that :- $v_p = \frac{\omega}{k}$, $\omega = k v_p$

$$v_g = \frac{d\omega}{dk} = \frac{d(k v_p)}{dk}$$

$$\begin{aligned} v_g &= v_p + k \frac{dv_p}{dk} \\ &= v_p + \frac{dv_p}{d\lambda} \times \frac{d\lambda}{dk} \end{aligned} \quad \left| \begin{array}{l} k = \frac{2\pi}{\lambda} \\ \frac{dk}{d\lambda} = -\frac{2\pi}{\lambda^2} \end{array} \right.$$

$$v_g = v_p - \frac{2\pi}{\lambda} \frac{dv_p}{d\lambda} \left(\frac{\lambda^2}{2\pi} \right)$$

$$\boxed{v_g = v_p - \lambda \frac{dv_p}{d\lambda}} \quad \text{--- (13)}$$

dispersion relation.

Case:- I v_p is independent of ' λ '

i.i. No dispersion at all

$$\therefore \frac{dv_p}{d\lambda} = 0$$

$$\therefore v_g = v_p \quad ! \quad \times \text{ reject.}$$

$\downarrow \qquad \downarrow$
 $v < c \quad \frac{c^2}{v} > c$

Case:- II $\frac{dv_p}{d\lambda} = -ve$ (dec. fm of λ)

$\lambda \uparrow \Rightarrow v_p \downarrow$ (\neq light) i.i. anomalous dispersion.

$$v_g = v_p + \underset{+ve}{(\text{Something})} \lambda$$

$$v_g > v_p, \text{ however } v_p > c$$

$$\therefore v_g > c \quad ! \quad \times \text{ reject.}$$

$$\text{but } \boxed{v_g = v} \quad \checkmark$$

Case:- III $\frac{dv_p}{d\lambda} = +ve$ (\uparrow fm. of λ)

$\lambda \uparrow \Rightarrow v_p \uparrow$ (= light) i.i. normal dispersion.

$$\therefore v_g = v_p - \underset{+ve}{(\text{Something})} \lambda$$

$$\Rightarrow v_g < v_p \quad \text{we need } v_g < c$$

we want

$$v_g < c$$

$$v_p - \lambda \frac{dv_p}{d\lambda} < c$$

or

$$v_p - c < \lambda \frac{dv_p}{d\lambda}$$

$$\left| \text{or } \frac{dv_p}{d\lambda} > \frac{v_p - c}{\lambda} \right|$$

de-Broglie wavelength of a charged particle:-

Relativistic case:-

q = charge, v = speed gained, V = accelerating potential.

m = mass in motion, p = linear momentum, m_0 = rest mass

$$\lambda = \frac{h}{mv} = \frac{h}{p} \quad \text{--- ①}$$

$$\text{Initial vel} = u = 0$$

$$W = \text{gain in K.E. (T)} \\ = E - u$$

$$\text{work done} = qV \quad \text{--- ②}$$

$$\text{Change in K.E.} = \text{final K.E.} - \text{initial K.E.}$$

$$= \text{Total energy} - \text{rest mass energy.}$$

$$= E - m_0 c^2$$

$$= \sqrt{m_0^2 c^4 + p^2 c^2} - m_0 c^2$$

③

work done = change in K.E.

$$qV = \sqrt{m_0^2 c^4 + p^2 c^2} - m_0 c^2$$

$$qV + m_0 c^2 = \sqrt{m_0^2 c^4 + p^2 c^2}$$

$$(qV + m_0 c^2)^2 = m_0^2 c^4 + p^2 c^2$$

$$q^2 V^2 + \cancel{m_0^2 c^4} + 2qV m_0 c^2 = \cancel{m_0^2 c^4} + p^2 c^2$$

$$p^2 c^2 = q^2 V^2 + 2qV m_0 c^2$$

$$p^2 = \frac{q^2 V^2}{c^2} + \frac{2qV m_0 c^2}{c^2}$$

$$p^2 = \frac{q^2 V^2}{c^2} \left(1 + \frac{2m_0 c^2}{qV} \right)$$

$$p = \frac{qV}{c} \sqrt{1 + \frac{2m_0 c^2}{qV}}$$

$$\lambda = \frac{h}{p}$$

$$\therefore \lambda = \frac{hc}{qV \sqrt{1 + \frac{2m_0 c^2}{qV}}}$$

→ applicable
for any particle
at any speed

★ Non-relativistic case (accelerating potential is low,

i.e. speed gained by the particle is less than the speed of light)

$$K.E. = \frac{1}{2} m v^2 = \frac{p^2}{2m}$$

$$\text{initial velocity} = u = 0$$

$$\therefore \text{change in K.E.} = \frac{1}{2} m v^2 - \frac{1}{2} m u^2$$

$$= \frac{p^2}{2m}$$

$$\text{work done} = qV$$

$$qV = \frac{p^2}{2m}$$

$$p^2 = 2m qV$$

$$p = \sqrt{2m qV}$$

$$\lambda = \frac{h}{\sqrt{2m qV}}$$

i.e. charged particle is e^- / positron.

$$q = 1.6 \times 10^{-19} \text{ C}$$

$$m = 9.1 \times 10^{-31} \text{ kg}$$

$$h = 6.6 \times 10^{-34} \text{ Js}$$

$$\text{or } \lambda = \sqrt{\frac{150}{V}} \text{ \AA}$$

(at low speed)

Particle:->

(i) It can't be omniscient. So a single de Broglie wave cannot be attached to a particle.

(ii) De Broglie wave travels faster than c , object less than c
 \therefore wave detach.

(iii) Probability of finding particle will be equal in the whole universe
 \therefore A single ^{de Broglie} wave cannot be attached to a particle.

$$\lambda = \frac{12.27 \times 10^{-10} \text{ m}}{\sqrt{V}}$$

$$\lambda = \frac{12.27 \text{ \AA}}{\sqrt{V}}$$

if External energy \approx rest mass energy.

Then speed $(v) \approx c$ ↓

10% of rest mass energy

$$= \frac{10}{100} \times m_0 c^2$$

$$= \frac{m_0 c^2}{10}$$

for external energy given $\leq 0.1 m_0 c^2 \Rightarrow$ (non-relativistic)

" " " " $> 0.1 m_0 c^2 \Rightarrow$ (relativistic)

for electron $m_0 c^2 = 511 \text{ kV}$

$$\frac{0.1 m_0 c^2}{1} = \frac{51.1 \text{ kV}}{1}$$

$= 51.1 \text{ kV} \approx$ accelerating potential.

$$\text{then } \lambda = \frac{12.27 \text{ \AA}}{\sqrt{U}}$$

for Proton, $m_0 c^2 = 931.5 \text{ MV}$

$$\frac{0.1 m_0 c^2}{1} = \frac{93.15 \text{ MV}}{1}$$

$= 93.15 \text{ MV} \approx$ accelerating potential.

$$\text{then } \lambda = \frac{h}{\sqrt{2m q U}}$$

if e^- beam is accelerated by 100 V

$$\therefore \lambda = \frac{12.27 \text{ \AA}}{\sqrt{100}}$$

$$\lambda = 1.227 \text{ \AA}$$

$$\frac{2 \times 10^{-19} = 1.9 \times 10^{-19}}{9.1 \times 10^{-31}}$$

$$4 \sqrt{2 \times 10^{-20+6+31}}$$

$$\frac{4}{3} \sqrt{\frac{2 \times 10^{-18}}{10}} \quad \begin{array}{r} 17 \\ 6 \\ 31 \\ 17 \end{array}$$

if e^- beam is accelerated by 5×10^6 V

$$\lambda = \frac{hc}{qV \sqrt{1 + \frac{2m_0c^2}{qV}}}$$

$$= \frac{6.63 \times 10^{-34} \times 3 \times 10^8}{1.6 \times 10^{-19} \times 10^6 \sqrt{1 + \frac{2 \times 9.1 \times 10^{-31} (3 \times 10^8)^2}{1.6 \times 10^{-19} \times 10^6}}}$$

$$V = pc \sqrt{\frac{1}{m_0^2 c^2 + p^2}}$$

Q1 - Broglie wavelength of a thermal particle:-

$$K.E. = \text{Thermal K.E.} = \frac{3}{2} kT$$

$$\frac{1}{2} m v^2 = \frac{3}{2} kT$$

$$\frac{p^2}{2m} = \frac{3}{2} kT$$

$$p^2 = 3 m kT$$

$$p = \sqrt{3 m kT}$$

$$\lambda = \frac{h}{\sqrt{3 m kT}}$$

$$T = \text{Kelvin}$$

$$k = 1.23 \times 10^{-38} \frac{\text{J}}{\text{K}}$$

$$m_e = 1.66 \times 10^{-27}$$

$$m_p = 1.840 m_e$$

* Electron $V = 11.24 \times 10^5 \text{ V}$

$$\lambda = \frac{12.27 \text{ \AA}}{\sqrt{11.24 \times 10^5}} = 1.15 \times 10^{-2} \text{ \AA} = 1.15 \text{ pm. (wrong answer)}$$

Energy given = qV
 $= 1 \times 11.24 \times 10^5 \text{ V}$
 $= 1124 \times 10^3 \text{ V}$
 $= 1124 \text{ kV} \gg 51.1 \text{ kV}$

$$K.E. = (m - m_0)c^2$$
$$= \left(\frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} - m_0 \right) c^2$$

$$qV = m_0 c^2 \left(\frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} - 1 \right)$$

$$1 \times 11.24 \times 10^5 = \left(\frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} - 1 \right) \times (511 \text{ keV})$$

$$v = 2.85 \times 10^8 \text{ m/sec.} < c$$

$$\lambda = \frac{hc}{qV \sqrt{1 + \frac{2m_0 c^2}{qV}}}$$

$$\lambda = 0.80 \text{ pm}$$



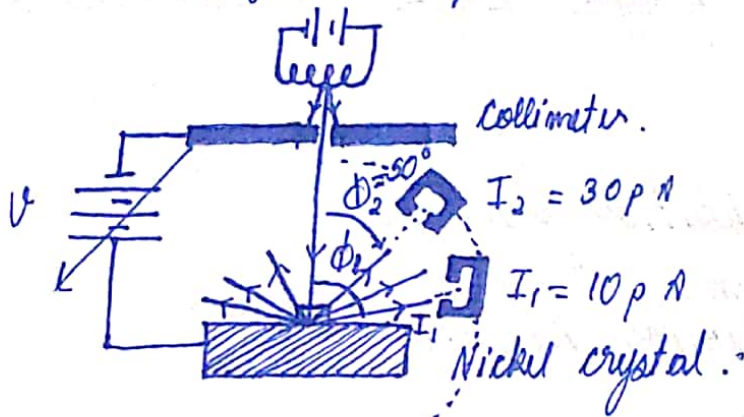
$$2\pi r = n\lambda$$

$$2\pi r = n \frac{h}{mv}$$

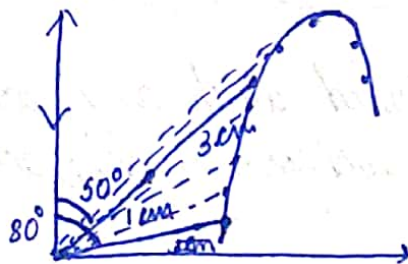
$$mvr = \frac{nh}{2\pi}$$

$$L = \frac{nh}{2\pi}$$

Davison and Germer experiment :->



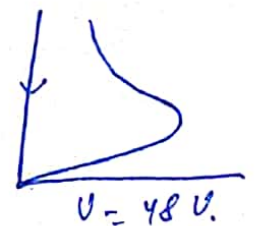
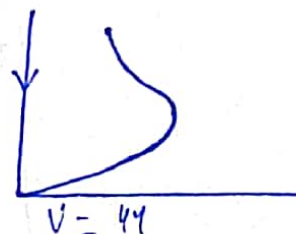
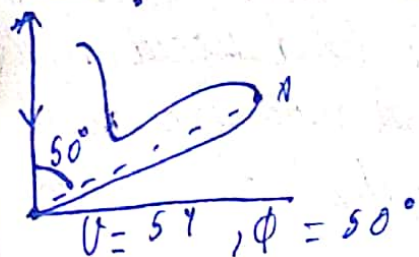
Incident beam :-

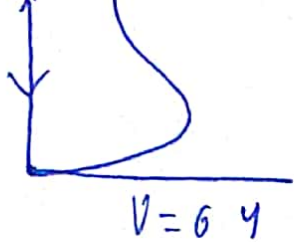


scale 1 cm = 10 pA

Bragg's law :- The diffraction from a crystal of ~~order~~ m^{th} order follows :-

$$2d \sin \theta = n\lambda$$





$$n = 1$$

$$d = 0.91 \text{ \AA}$$

$$\theta = 65^\circ$$

$$2 \times 0.91 \sin 65^\circ = 1 \times \lambda$$

$$\lambda = 1.65 \text{ \AA}$$

$$\lambda = \frac{12.27}{\sqrt{V}}$$

$$= \frac{12.27}{\sqrt{54}} = 1.67 \text{ \AA}$$

Comparison b/w de-Broglie waves and Electromagnetic waves:-

I. Similarities:-


1. Both waves show superposition principle.
2. Can pass through vacuum.
3. Both show normal dispersion (i.e. $\lambda \uparrow \Rightarrow v_p \uparrow$)

II. Differences:-

Electromagnetic waves :->

1. associated only with accelerating charged particle
2. Not deflected by electric & magnetic field

De-Broglie waves :->

1. associated with any particle at rest, uniform motion, accelerating.
2.  The wave packets always associated with the particle and the particle is always associated with wave group.
 \therefore The de-Broglie waves may or may not be deflected by the Electric and magnetic field depending on the nature of the particle.

3. Travels in vacuum with speed 'c' $3 \times 10^8 \text{ m/sec}$
4. Do not show dispersion in vacuum or show dispersion in medium only or all EM waves of diff λ/ν travels with same speed in vacuum

$$3. v_p = \frac{c^2}{v} > c$$

$$v_g = v < c$$

4. They show dispersion in vacuum as well as in medium.

Basics for Schrodinger

$\psi \rightarrow$ calculatable

\rightarrow not measurable/observable

\rightarrow has no significance in itself.

$$\boxed{\frac{d^2 \psi}{dx^2} + \frac{2m}{\hbar^2} (\epsilon - u) \psi = 0} \rightarrow \text{1-D schrodinger eqn.}$$

or

$$\boxed{\frac{d^2 \psi}{dx^2} + \frac{8\pi^2 m}{h^2} (\epsilon - u) \psi = 0}$$

|||

$$\boxed{\frac{d^2 y}{dt^2} + \omega^2 y = 0}$$

$$\boxed{\psi = y, \quad x = t, \quad \frac{8\pi^2 m}{h^2} (\epsilon - u) = \omega^2}$$

soln.

$$\boxed{y = A \sin \omega t + B \cos \omega t}$$

$$\boxed{\omega = \sqrt{\frac{K}{m}}}$$

$$\therefore \Psi = f(x, t)$$

for 3-D $\rightarrow \Psi = f(x, y, z, t) \therefore$ wave fn. is calculatable

In 1-D $\rightarrow \Psi = \frac{1}{\sqrt{\text{length}}} \rightarrow$ so it cannot be measured

Although Ψ has no significance and it's not even measurable then why it is useful?

Operators in quantum mechanics \rightarrow

\downarrow
Any process which changes one fn. to another fn.

let \hat{A} is an operator.

$\hat{A} = \sqrt{\quad}$ \rightarrow no significance.
= square root of -----

$$\hat{A}(2) = \sqrt{2} = 1.414$$

$$\hat{A}^2(4) = \sqrt{\sqrt{4}}$$

$$\hat{A}(3) = \sqrt{3} = 1.732$$

$$= \sqrt{2} = 1.414$$

$$\hat{A}(4) = \sqrt{4} = 2$$

$$\hat{A}(\Psi) = \sqrt{\Psi}$$

$$\hat{A}(4x^2) = \sqrt{4x^2} = 2x$$

let $\hat{A} = \frac{d}{dx}$ = diff. of ----- w.r.t. 'x'

$$\Psi = \sin 4x$$

$$\therefore \hat{A}(\psi) = \frac{d(\psi)}{dx} = \text{diff. of } \sin^4 x \text{ w.r.t. } x$$

$$= \frac{d(\sin^4 x)}{dx} = 4 \cos^4 x$$

$$\text{Let } \hat{B} = \frac{d^2}{dx^2} = \left(\frac{d}{dx}\right) \left(\frac{d}{dx}\right)$$

$$= (\hat{A})(\hat{A}) = (\hat{A})^2$$

$$\therefore \hat{B}(\psi) = \hat{A}^2(\psi) = \frac{d^2}{dx^2}(\psi) = \frac{d^2}{dx^2}(\sin^4 x)$$

$$= -16 \sin^4 x$$

• Linear operator :-

$$\boxed{\hat{A}(C_1 \psi_1 + C_2 \psi_2) = C_1 \hat{A} \psi_1 + C_2 \hat{A} \psi_2}$$

$$\text{Let, } \hat{A} = \frac{d}{dx}$$

all diff. operators.
"integral" .

$$\frac{d}{dx}(\psi_1 \pm \psi_2) = \frac{d\psi_1}{dx} \pm \frac{d\psi_2}{dx}$$

$$\hat{A} = \sqrt{\quad}$$

$$\hat{A}(4) = \sqrt{4} = 2$$

$$\hat{A}(9) = \sqrt{9} = 3$$

$$\hat{A}(4+9) = \sqrt{13} = 3.5 \text{ sm.}$$

i.e. $\hat{A}(4+9) \neq \hat{A}(4) + \hat{A}(9)$
 \therefore (non linear fm)

The operators that we use in quantum mechanics are linear operators:-

* Eigen value eqn.:->

- Eigen fm.
- Eigen operator.
- Eigen value

Definition:-

$$\hat{A}(\psi) = \alpha \psi$$

Eigen value eqn.
Eigen fm

↑
Eigen Operator.

Const. = real or imaginary.
(Secular eqn.)

example:-> let $\hat{A} = \frac{d^2}{dx^2}$, $\psi = \sin 4x$

$$\hat{A}(\psi) = -16 \sin 4x$$

$$\boxed{\hat{A}(\psi) = -16 \psi}$$

↓ ↓ ↓
 $\frac{d^2}{dx^2}$ Const. Eigen fm.

$$\hat{A} = \frac{d}{dx}, \quad \psi = e^{4x}$$

$$\boxed{\hat{A}(\psi) = \frac{d(e^{4x})}{dx} = 4e^{4x} = 4\psi}$$

$$\psi = \sin 4x, \quad \hat{A} = \frac{d}{dx}$$

$$\hat{A}(\psi) = \frac{d(\sin 4x)}{dx} = 4 \cos 4x$$

$$\boxed{\hat{A}(\psi) = 4 \psi'}$$

The quantum mechanical operators are always fixed, so, we have to choose an appropriate wave fm so, that the eigen value eqn. holds.

$$\boxed{\hat{A}(\psi) = \tilde{\alpha} \psi} \text{ Matrix eqn.}$$

$$\text{Let } \psi = A e^{-i(\omega t - \vec{k} \cdot \vec{r})} \quad \text{--- (1)}$$

$$\vec{r} = x\hat{i} + y\hat{j} + z\hat{k}$$

$$\vec{k} = k_x\hat{i} + k_y\hat{j} + k_z\hat{k}$$

$$\vec{k} \cdot \vec{r} = k_x x + k_y y + k_z z$$

$$\hat{p}_x = -i\hbar \frac{\partial}{\partial x}$$

$$\hat{p}_x(\psi) = -i\hbar \frac{\partial \psi}{\partial x}$$

$$= -i\hbar \frac{\partial}{\partial x} \{ A e^{-i(\omega t - \vec{k} \cdot \vec{r})} \}$$

$$= [i\hbar A e^{-i(\omega t - \vec{k} \cdot \vec{r})}] \times \frac{\partial}{\partial x} [-i(\omega t - \vec{k} \cdot \vec{r})]$$

$$= [i\hbar \psi] (-i) [0 - (k_x(1) + 0 + 0)]$$

$$= +k_x \hbar \psi$$

$$\begin{aligned} \because p &= \hbar k \\ p_x &= \hbar k_x \end{aligned}$$

$$\boxed{\hat{p}_x(\psi) = p_x \psi} \rightarrow \text{Eigen value eqn.}$$

\downarrow operator. \downarrow Const. \downarrow Eigen fn.

$$\hat{p}_y = -i\hbar \frac{\partial}{\partial y}$$

$$\therefore \boxed{\hat{p}_y(\psi) = p_y \psi}$$

similarly.

$$\boxed{\hat{p}_z(\psi) = p_z \psi}$$

$$\hat{p} = -i\hbar \vec{\nabla}$$

$$\hat{p} = \hat{i}(-i\hbar \frac{\partial}{\partial x}) + \hat{j}(-i\hbar \frac{\partial}{\partial y}) + \hat{k}(-i\hbar \frac{\partial}{\partial z})$$

$$\hat{p} = \hat{i}\hat{p}_x + \hat{j}\hat{p}_y + \hat{k}\hat{p}_z$$

$$-i\hbar k_x \psi$$

$$p_x \psi$$

$$\hat{p}(\psi) = \hat{i}p_x(\psi) + \hat{j}p_y(\psi) + \hat{k}p_z(\psi)$$

$$\hat{p}(\psi) = (ip_x + \hat{j}p_y + \hat{k}p_z) \psi$$

$$\hat{p}(\psi) = \vec{p} \psi$$

★ Hamiltonian operator: $\rightarrow (\hat{H}) \rightarrow$ (time independent form.)

$$\hat{H} = i\hbar \frac{\partial}{\partial t}$$

time dependent form.

$$\hat{H}(\psi) = i\hbar \frac{\partial}{\partial t} e^{-i(\omega t - \vec{k} \cdot \vec{r})}$$

$$\hat{H}(\psi) = i\hbar [e^{-i(\omega t - \vec{k} \cdot \vec{r})}] (-i\omega)$$

$$\hat{H}(\psi) = \hbar \omega \psi \quad \therefore \hat{H} \psi = E \psi$$

time independent:

$$\hat{H}(\psi) = \frac{-\hbar^2}{2m} \nabla^2(\psi) + U(\psi)$$

$$\hat{H} = \frac{-\hbar^2}{2m} \nabla^2 + U$$

$$\nabla^2 = \vec{\nabla} \cdot \vec{\nabla} = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}$$

$$\nabla^2(\psi) = \frac{\partial^2 \psi}{\partial x^2} + \frac{\partial^2 \psi}{\partial y^2} + \frac{\partial^2 \psi}{\partial z^2}$$

$$\frac{\partial \psi}{\partial x} = \frac{\partial}{\partial x} \{ A e^{-i(\omega t - \vec{k} \cdot \vec{r})} \}$$

$$= \{ A e^{-i(\omega t - \vec{k} \cdot \vec{r})} \} \frac{\partial}{\partial x} (-i(\omega t - \vec{k} \cdot \vec{r}))$$

$$= -i\psi \{ (0 - (k_x(1) + 0 + 0)) \}$$

$$= i k_x \psi$$

$$\frac{\partial \psi}{\partial x} = i k_x A e^{-i(\omega t - \vec{k} \cdot \vec{r})}$$

$$\frac{\partial^2 \psi}{\partial x^2} = i k_x \frac{\partial \psi}{\partial x}$$

$$= i(k_x)(i k_x \psi)$$

$$\boxed{\frac{\partial^2 \psi}{\partial x^2} = -k_x^2 \psi}$$

$$\therefore \boxed{\frac{\partial^2 \psi}{\partial y^2} = -k_y^2 \psi}$$

$$\therefore \frac{\partial^2 \psi}{\partial z^2} = -k_z^2 \psi$$

$$\nabla^2 \psi = -(k_x^2 + k_y^2 + k_z^2) \psi$$

$$= -k^2 \psi$$

$$\therefore \hat{H}(\psi) = \frac{-\hbar^2}{2m} (-k^2 \psi) + U(\psi)$$

$$= \frac{\hbar^2 k^2}{2m} \psi + U(\psi)$$

$$= \frac{p^2}{2m} \psi + U\psi$$

$$\boxed{p = \hbar k}$$

$$\hat{H}(\psi) = \left(\frac{p^2}{2m} + U \right) \psi$$

$$\boxed{K \cdot E + P \cdot E = E}$$

$$\therefore \boxed{\hat{H}(\psi) = E \psi}$$

→ The wave fm. is also called state fm. because wave fm. gives all physical parameter of the physical state.

Position operator:-

$$\begin{cases} \hat{x} = x \\ \hat{y} = y \\ \hat{z} = z \end{cases}$$

$$\boxed{\hat{R} = (x\hat{i} + y\hat{j} + z\hat{k})}$$

Potential energy operation:-

$$\boxed{\hat{U} = U}$$

(never depends upon velocity, momentum for 1-D harmonic oscillator. -1 um)

$$\boxed{\hat{U} = \frac{1}{2} k x^2}$$

or
any derivative $\frac{d}{dt}$

Kinetic energy operator:-

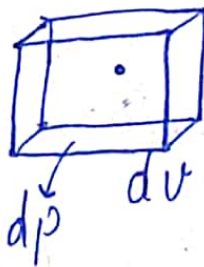
$$\hat{T} = \frac{-\hbar^2}{2m} \nabla^2$$

$$E = \frac{p^2}{2m} + U$$

$$\frac{dE}{dp} = \frac{p}{m} + 0$$

Normalization of a wave fm.:-

$$\boxed{|\psi|^2 \propto \frac{dP}{dV}}$$



$\frac{dP}{dV}$ = Probability density.

$$\Rightarrow \boxed{|\psi|^2 = N \frac{dP}{dV}} \quad \text{--- } \textcircled{1}$$

If $N = 1$
 $\Rightarrow |\psi|^2 = \frac{dP}{dV}$

i.e. $|\psi|^2 \rightarrow$ exactly equal to p probability density.

Then \rightarrow

$\psi \rightarrow$ Normalized w.f.

if $|\psi|^2 \neq \frac{dP}{dV}$ or $N \neq 1$

$$\frac{dP}{dV} \propto |\psi|^2$$
$$\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} dP = N |\psi|^2 dV = 1$$

Then \rightarrow $\psi \rightarrow$ Unnormalized w.f.

It is always possible to obtain normalized w.f. from unnormalized w.f. by multiplying ψ with some suitable const..

w.f. = $\psi \rightarrow$ soln of Schrodinger eq. (S.E.)

S.E. \rightarrow differential eq.

That if ψ is a soln. of a differential eq. then $c\psi$ is also a soln. of diff. eq.

Let ψ be the soln. obtained by solving a S.E.

Let $|\psi|^2 \rightarrow$ known

Find $\int_{-\infty}^{\infty} |\psi|^2 dV$

Let $\int_{-\infty}^{\infty} |\psi|^2 dV = N$

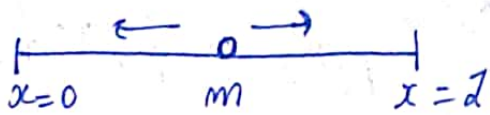
$$\Rightarrow \int_{-\infty}^{\infty} \frac{|\psi|^2}{N} dV = 1$$

Put $\frac{\psi}{\sqrt{N}} = \psi'$

$$\therefore \int_{-\infty}^{\infty} |\psi|^2 dV = 1$$

$$\therefore \boxed{|\psi|^2 = \frac{dP}{dV}}$$

$\Psi' = \frac{\Psi}{\sqrt{N}}$ is normalized



$$\Psi = A \sin\left(\frac{n\pi x}{L}\right) \quad (\text{given})$$

We assert that A is chosen in such a way that

$$|\Psi|^2 = \frac{d\rho}{dx} \quad (\text{exactly})$$

$$\Rightarrow d\rho = |\Psi|^2 dx$$

$$\int_0^L d\rho = 1$$

$$\rightarrow \int_0^L |\Psi|^2 dx$$

$$\rightarrow \int_0^L A^2 \sin^2\left(\frac{n\pi x}{L}\right) dx = 1$$

$$\frac{A^2}{2} \int_0^L \left[1 - \cos\left(\frac{2n\pi x}{L}\right)\right] dx$$

$$\frac{A^2}{2} [L - 0] = 1$$

$$\therefore A = \sqrt{\frac{2}{L}}$$

$$\Psi = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right)$$

Hermitian Operation: \rightarrow

$$\int_{-\infty}^{\infty} \psi_m^* \hat{A} \psi_m dV = \int_{-\infty}^{\infty} \psi_m (\hat{A} \psi_m)^* dV \quad \text{--- (1)}$$

Importance: \rightarrow

Let \hat{A} is Hermitian operator.

$\therefore \hat{A}$ satisfies eqn --- (1)

Now $\psi_m = \psi_m$ (say)

$$\int_{-\infty}^{\infty} \psi_m^* \hat{A} \psi_m dV = \int_{-\infty}^{\infty} \psi_m (\hat{A} \psi_m)^* dV$$

$$\Rightarrow \int_{-\infty}^{\infty} \psi_m^* \hat{A} \psi_m dV = \int_{-\infty}^{\infty} \psi_m (\hat{A} \psi_m)^* dV$$

Let $\hat{A}(\psi_m) = \alpha_m \psi_m$

$$\Rightarrow \int_{-\infty}^{\infty} \psi_m^* \alpha_m \psi_m dV = \int_{-\infty}^{\infty} \psi_m \psi_m^* \alpha_m^* dV$$

$$\alpha_m \int_{-\infty}^{\infty} \psi_m^* \psi_m dV = \alpha_m^* \int_{-\infty}^{\infty} \psi_m \psi_m^* dV$$

$$\alpha_m \int_{-\infty}^{\infty} |\psi_m|^2 dV = \alpha_m^* \int_{-\infty}^{\infty} |\psi_m|^2 dV$$

$$\therefore (\alpha_m - \alpha_m^*) = 0 \text{ or } \int_{-\infty}^{\infty} |\psi_m|^2 dV = 0 \text{ (rejected)}$$

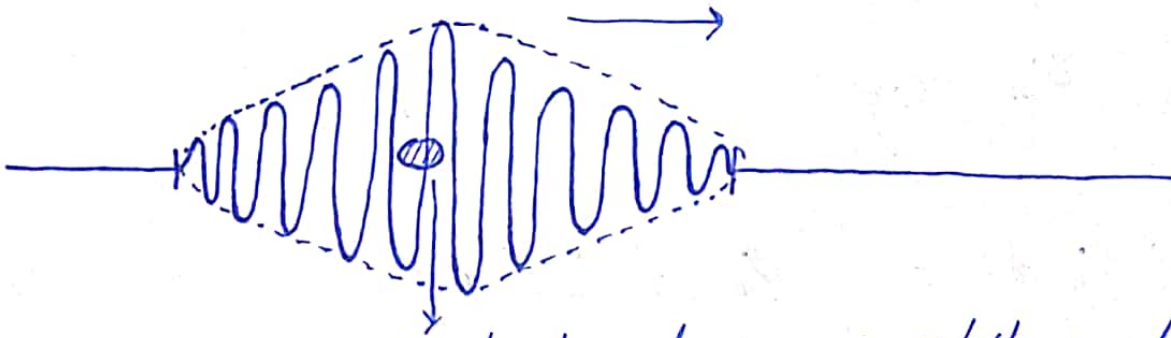
$$\boxed{\alpha_m = \alpha_m^*} \therefore \alpha_m = \text{a real no.}$$

$\frac{d}{dt} = \text{explicit}$
for

$\frac{d}{dt} = \text{implicit}$
for

Eigen values of Hermitian operators are always real.

Wave mechanical Treatment :->



Can ~~the~~ travel in any arbitrary direction inside the wave packet.

The wave mechanical treatment is that method in which we study the average motion of a particle by observing the motion of wave packet attached with it.

Newtonian

x, v or p directly measurable with 100% accuracy

(deterministic approach)

wave mechanical:

we study the motion of wave packet (gives us the average motion of the particle)

(undeterministic approach / or (Probabilistic))

Schrödinger Equation:- (eqn. of motion of a wave packet attached with a particle)

Time independent SE (TISE)

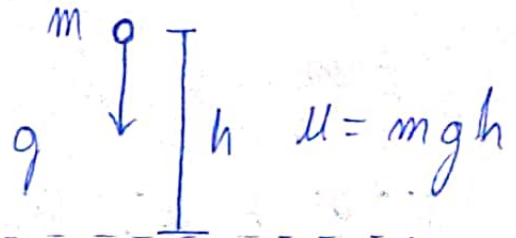
Time dependent SE (TDSE)

$$\Psi(\vec{x}, \vec{p}, t) = \psi(\vec{x}, \vec{p}) \phi(t)$$

$$\Psi(\vec{x}, \vec{p}, t) \neq \psi(\vec{x}, \vec{p}) \phi(t)$$

is possible when potential energy 'U' does not depend upon 't' explicitly.

example (Josephson effect)
Potential energy depends explicitly on 't'



'U' is \downarrow with $g \propto t$

$$U = f(h)$$

$$h = f(t)$$

$\therefore U \propto \frac{1}{t}$
depends on 't' implicitly through 'h'

TISE:-

$$\psi_m = A e^{-i(\omega_m t - \vec{k}_m \cdot \vec{r}_m)} \quad \text{--- ①}$$

$\omega = \sqrt{-1}$, ω = angular freq.
 k = propagation const.

$$v_p = \frac{\omega}{k} \quad \text{--- ②}$$

Ψ must satisfy standard wave eqn.

$$\nabla^2 \psi - \frac{1}{v_p^2} \frac{\partial^2 \psi}{\partial t^2} = 0 \quad \text{--- ③}$$

h

$$\frac{\partial \psi}{\partial t} = \left[A (-i\omega + -k \cdot \nabla) \right] (-i\omega - 0)$$

$$\boxed{\frac{\partial \psi}{\partial t} = -i\omega \psi} \quad (4)$$

$$\frac{\partial^2 \psi}{\partial t^2} = (-i\omega) \left(\frac{\partial \psi}{\partial t} \right)$$

$$\frac{\partial^2 \psi}{\partial t^2} = -i\omega (-i\omega \psi)$$

$$\boxed{\frac{\partial^2 \psi}{\partial t^2} = -\omega^2 \psi} \quad (5)$$

from (5) & (3)

$$\nabla^2 \psi + \frac{\omega^2}{V_p^2} \psi = 0$$

$$\boxed{\nabla^2 \psi + k^2 \psi = 0} \quad (6)$$

Dirac eqn.

Assume :- particle is travelling non-relativistically.

$$K.E. \approx \frac{1}{2} m v^2 = \frac{p^2}{2m}$$

$$= \frac{\hbar^2 k^2}{2m}$$

$$E = K \cdot E + P \cdot E.$$

$$= \frac{\hbar^2 k^2}{2m} + U$$

$$E - U = \frac{\hbar^2 k^2}{2m}$$

$$k^2 = \frac{2m}{\hbar^2} (E - U) \quad \text{--- (1)}$$

$$\nabla^2 \psi + \frac{2m}{\hbar^2} (E - U) \psi = 0$$

3-D TISE for a restricted particle / particle subjected to some ext. force.

for a free particle $U = 0$

$$\therefore \nabla^2 \psi + \frac{2mE}{\hbar^2} \psi = 0$$

for 1-D

$$\nabla^2 \psi \longrightarrow \frac{d^2 \psi}{dx^2}$$

TISE:-

$$\psi = A e^{-i(\omega t - \vec{k} \cdot \vec{r})}$$

Non relativistic motion is assumed.

$$E = K \cdot E + P \cdot E.$$

$$E = \frac{p^2}{2m} + U$$

multiply by ψ

$$E \psi = \frac{p^2}{2m} \psi + U \psi \Rightarrow E \psi = \frac{\hbar^2 k^2}{2m} \psi + U \psi$$

Take partial derivative of ψ w.r.t. t

$$\frac{\partial \psi}{\partial t} = \left[A e^{-i(\omega t - \vec{k} \cdot \vec{r})} \right] (-i\omega)$$
$$= \psi (-i\omega)$$

multiply both sides by $i\hbar$

$$i\hbar \frac{\partial \psi}{\partial t} = i\hbar (\psi) (-i\omega)$$

$$i\hbar \frac{\partial \psi}{\partial t} = \hbar \omega \psi$$

$$\therefore i\hbar \frac{\partial \psi}{\partial t} = E \psi \quad \text{--- (3)}$$

Take partial derivative of ψ w.r.t. x

$$\frac{\partial \psi}{\partial x} = \left[A e^{-i(\omega t - \vec{k} \cdot \vec{r})} \right] [i\hbar k_x]$$

$$\frac{\partial \psi}{\partial x} = i\hbar k_x \psi \quad \text{--- (4)}$$

Similarly :-

$$\frac{\partial^2 \psi}{\partial y^2} = i k_y \psi, \quad \frac{\partial^2 \psi}{\partial z^2} = i k_z \psi$$

$$\frac{\partial^2 \psi}{\partial x^2} = i k_x \frac{\partial \psi}{\partial x} = i k_x (i k_x \psi)$$

$$\therefore \frac{\partial^2 \psi}{\partial x^2} = -k_x^2 \psi \quad \text{--- (5)}$$

Similarly :-

$$\frac{\partial^2 \psi}{\partial y^2} = -k_y^2 \psi \quad \text{--- (6)} \quad , \quad \frac{\partial^2 \psi}{\partial z^2} = -k_z^2 \psi \quad \text{--- (7)}$$

$$\therefore \nabla^2 \psi = -k^2 \psi \quad \text{--- (8)}$$

$$\therefore E \psi = \frac{\hbar^2 k^2}{2m} \psi + U \psi$$

$$i \hbar \frac{\partial \psi}{\partial t} = \frac{-\hbar^2 \nabla^2 \psi}{2m} + U \psi \quad \text{***}$$

3-D TISE for a restricted particle

for a free particle $U = 0$

$$i \hbar \frac{\partial \psi}{\partial t} = \frac{-\hbar^2 \nabla^2 \psi}{2m}$$

for 1-D

$$\nabla^2 \psi \rightarrow \frac{\partial^2 \psi}{\partial x^2}$$

Limitations of Schrödinger eqn. :-

① Applicable only for non relativistic problems.

② 1st particle. 2nd particle - - - - -
(x_1, y_1, z_1), (x_2, y_2, z_2), - - - - -

$$\nabla^2 \Psi + \frac{2m}{\hbar^2} (E - U) \Psi = 0$$

for 1st particle eqn :- $\nabla_1^2 \Psi_1 + \frac{2m_1}{\hbar^2} (E_1 - U_1) \Psi_1 = 0$

for 2nd particle eqn :- $\nabla_2^2 \Psi_2 + \frac{2m_2}{\hbar^2} (E_2 - U_2) \Psi_2 = 0$

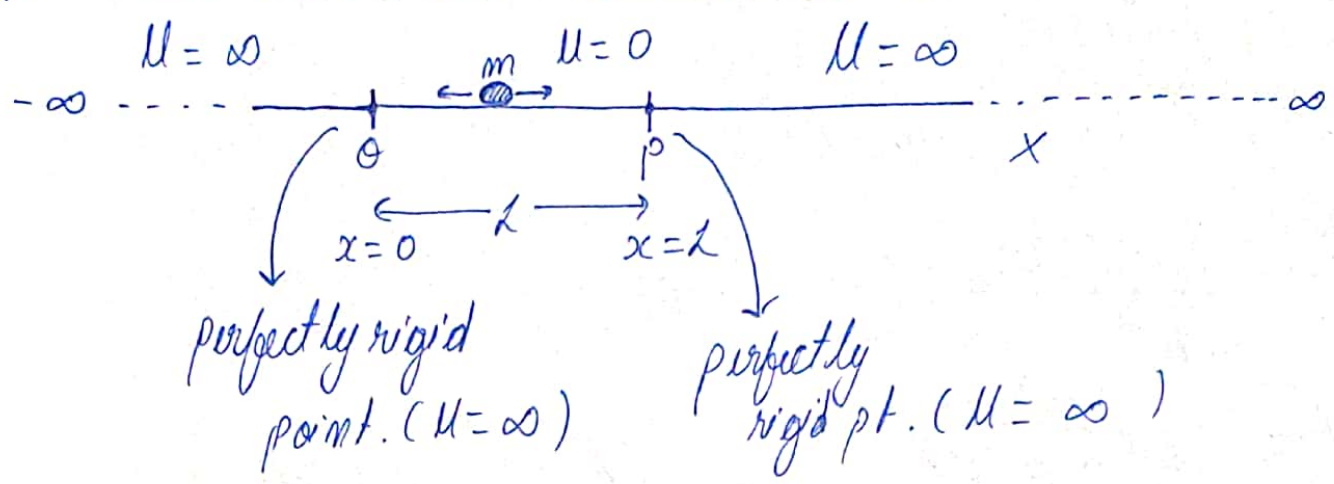
i.e. for many particle system this ~~is~~ ~~to~~ Schrödinger eqn. is very difficult to solve

③ for solving Schrödinger eqn. we require very high performance computational tools

④ Schrödinger eqn. on solving gives us many solutions out of which large no. of solutions are meaningless

⑤ Schrödinger eqn. tells us ~~also~~ only about n, l and m but does not provide any result for 's' (spin of no.)

Particle in a one dimensional box :->

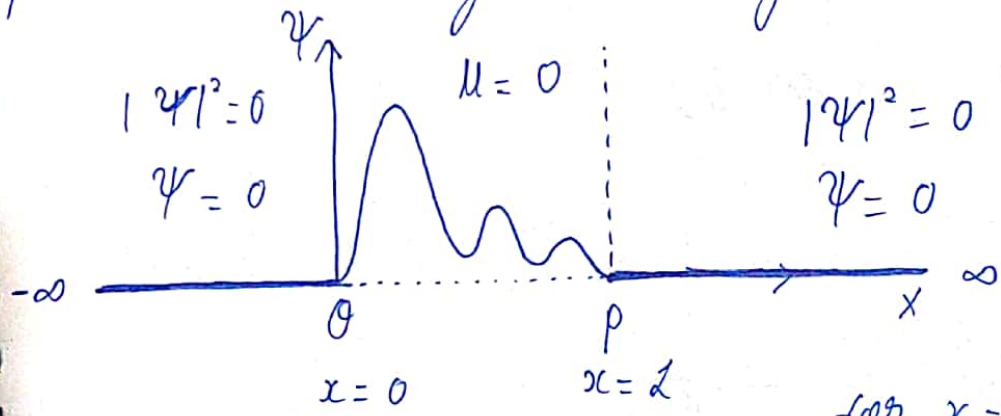


We know that every system tends to occupy the state of minimum potential energy

since, the particle is unable to cross 'o' and 'p' pt.

\therefore In quantum mechanical way we can say that perfectly rigid pt. or perfectly rigid region must have infinite potential energy

Infinite potential energy gives instability to particle so, particle will never go in that region



$$\left. \begin{array}{l} \text{for } x=0, \psi=0 \\ \text{for } x=l, \psi=0 \end{array} \right\} \text{--- (1)}$$

Assume b/w 'o' and 'p' particle is not interacting with anything.
 b/w 'o' and 'p' $U = 0$

\therefore Total energy from o to p is kinetic energy

1-D TISE in region $\psi \rightarrow$

$$\frac{d^2\psi}{dx^2} + \frac{2m(E-U)\psi}{\hbar^2} = 0$$

(But $U=0$)

$$\therefore \frac{d^2\psi}{dx^2} + \frac{2mE}{\hbar^2} \psi = 0 \quad \text{--- (2)}$$

Suppose $\frac{2mE}{\hbar^2} = k^2$ --- (3) (k is any const.)

$$\therefore \frac{d^2\psi}{dx^2} + k^2 \psi = 0 \quad \text{--- (4)}$$

soln. of (4) is of the form.

$$\psi = A \sin(kx) + B \cos(kx) \quad \text{--- (5)}$$

for $x=0$, $\psi=0$

$$\therefore 0 = A \sin(0) + B \cos(0)$$

$$\therefore B=0$$

$$\therefore \psi = A \sin(kx) \quad \text{--- (6)}$$

for $x=L$, $\psi=0$

$$0 = A \sin(kL) \quad \text{--- (7)}$$

either $A=0$ or $\sin(kL)=0$

(non logical soln.)

$$a = -\omega^2 y$$
$$\frac{d^2 y}{dt^2} + \omega^2 y = 0$$

$$y = A \sin(\omega t) + B \cos(\omega t)$$

$$\therefore \text{if } A = 0$$

$$\Psi = 0 \text{ (inside)}$$

$$\Psi = 0 \text{ (outside)}$$

\therefore particle is not present in the whole universe.

$$\sin(kL) = 0$$

$$\therefore \boxed{kL = m\pi} \text{ --- } \textcircled{8} \text{ (quantum condition)}$$

(stability ")

$$\text{for } m = 0$$

$$A \neq 0$$

$$\therefore kL = 0$$

$$\Psi = A \sin\left(\frac{m\pi}{L}x\right) \text{ (rejected)}$$

$$\therefore \Psi = 0$$

$$m = -1$$

$$kL = -\pi$$

\therefore (rejected)

$m = -ve$ (rejected)

$$\frac{(+ve)(+ve)}{\lambda}$$

$$\frac{2\pi}{\lambda}$$

$$\therefore m = 1, 2, 3, \dots$$

$$k = \frac{m\pi}{L}, \quad k^2 = \frac{m^2\pi^2}{L^2}$$

$$\therefore \frac{2mE}{\hbar^2} = \frac{m^2\pi^2}{L^2}$$

$$E = \frac{m^2\pi^2\hbar^2}{2mL^2}$$

$$E = m^2 (\mathcal{Q})$$

$$\therefore |E_m = m^2 (\mathcal{Q})| \text{ } \star \star \star$$

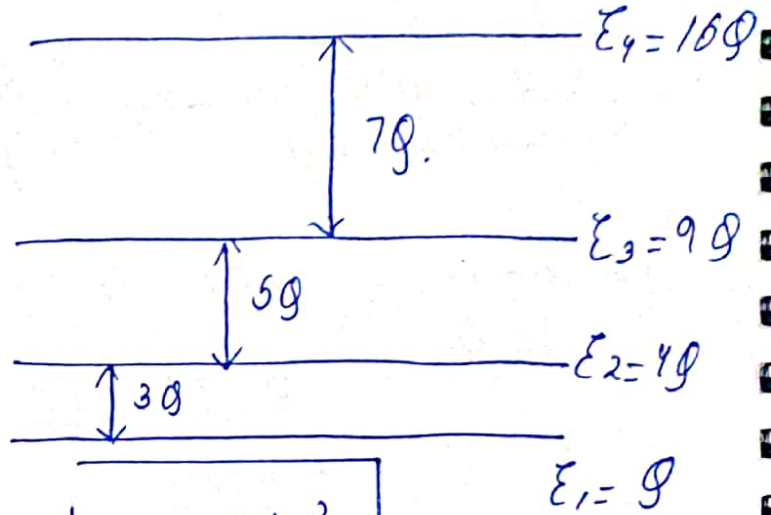
$$\left[\mathcal{Q} = \frac{\pi^2 \hbar^2}{2mL^2} \right]$$

$$\epsilon_1 = 1^2 \phi = \phi \text{ (ground state energy)}$$

$$\epsilon_2 = 2^2 \phi = 4\phi \text{ (1st excited state)}$$

$$\epsilon_3 = 3^2 \phi = 9\phi \text{ (2nd " ")}$$

$$\epsilon_4 = 4^2 \phi = 16\phi \text{ (3rd " ")}$$



$$\phi = \frac{\pi^2 \hbar^2}{2mL^2}$$

$$\mu = 0, \therefore \boxed{E_n = K \cdot E}^{n \text{th}}$$

$$\therefore E_{\text{min}} = K \cdot E_{\text{min}} = \epsilon_1 = \phi \neq 0$$



never at rest.

$$\frac{A^2}{2} [2 - 0] = 1\phi$$

$$\therefore A = \sqrt{\frac{2}{L}}$$

$$\Psi = A \sin(kx)$$

$$\Psi = A \sin\left(\frac{n\pi x}{L}\right) \quad \text{--- (10)}$$

Assertion:- 'A' is chosen in such a way that Ψ becomes normalized.

i.e. $|\Psi|^2 =$ exactly probability density

$$= \frac{dP}{dx}$$

$$\Rightarrow \int_{x=0}^{x=L} dP = 1$$

$$\int_{x=0}^{x=L} \phi |\Psi|^2 dx = 1$$

$$A^2 \int_0^L \sin^2\left(\frac{n\pi x}{L}\right) dx = 1\phi$$

$$\frac{A^2}{2} \int_0^L \left(1 - \cos\left(\frac{2n\pi x}{L}\right)\right) dx = 1$$

∴ Normalized wave fn.

$$\Psi = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi}{L} x\right) \quad \text{--- (11)}$$

→ Eigen fn.

$$E_n = n^2 \phi = \frac{n^2 \pi^2 \hbar^2}{2mL^2} \quad \text{--- (12)}$$

→ Energy eigen value fn.

$$E_n = \phi, 4\phi, 9\phi, 16\phi, 25\phi, 36\phi, \dots$$

Why not $1.5\phi, 2\phi, 3\phi \dots$? (not permissible levels)

$$\left[\because L = \frac{nh}{2\pi} = n\hbar \right]$$

$$k = \frac{n\pi}{L} \quad , \quad p = \hbar k$$

$$p = \hbar \left(\frac{n\pi}{L} \right)$$

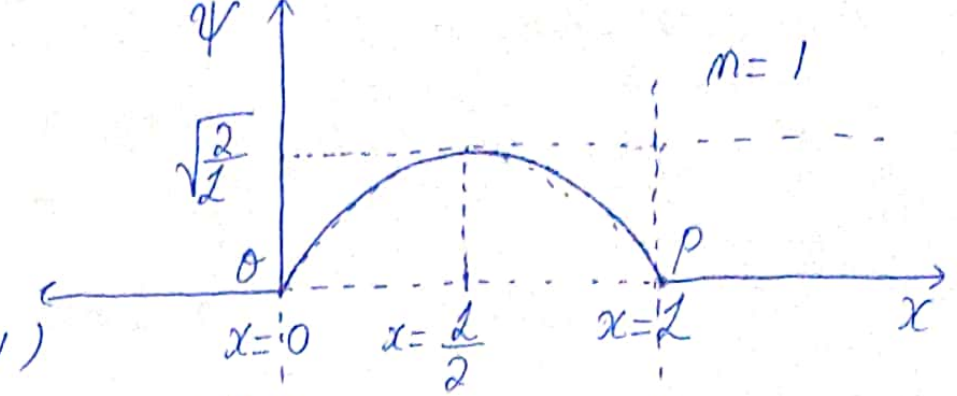
$$p = \frac{h}{2\pi} \left(\frac{n\pi}{L} \right)$$

$$p = \frac{nh}{2L} \quad \text{--- (13)} \quad \left| \quad L = \frac{nh}{2\pi} \right.$$

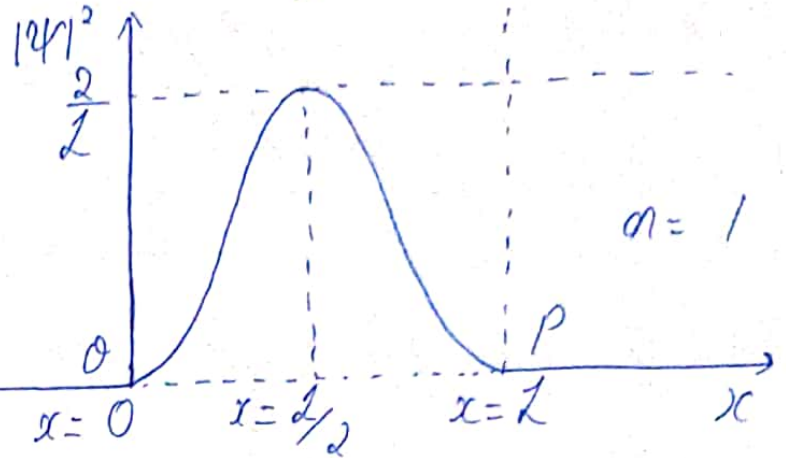
$$\psi = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right)$$

$$\psi = \sqrt{\frac{2}{L}} \sin\left(\frac{\pi x}{L}\right)$$

(for $n=1$)



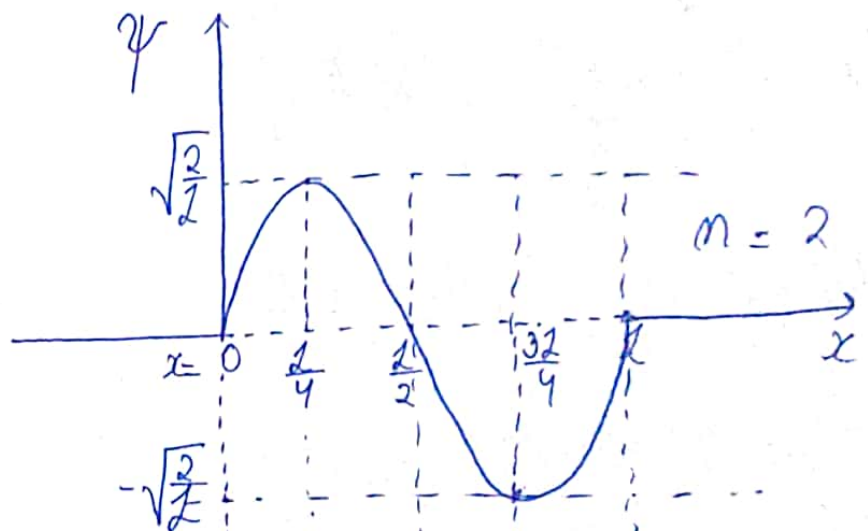
$x=0$	$x=\frac{L}{2}$	$x=L$
$\psi=0$	$\psi=\sqrt{\frac{2}{L}}$	$\psi=0$



for $n=2$

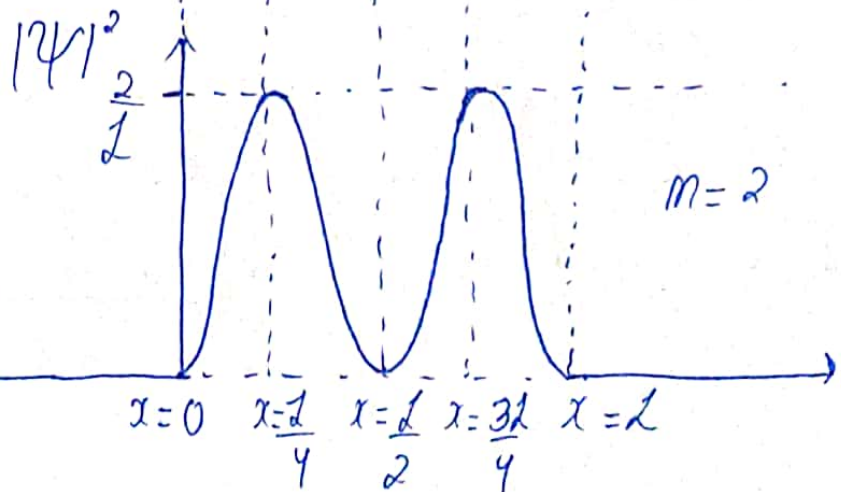
$$\psi = \sqrt{\frac{2}{L}} \sin\left(\frac{2\pi x}{L}\right)$$

$x=0$	$x=\frac{L}{2}$	$x=L$
$\psi=0$	$\psi=\sqrt{\frac{2}{L}}$	$\psi=0$



$x=\frac{3L}{4}$	$x=L$
$\psi=\sqrt{\frac{2}{L}}$	$\psi=0$

$x=\frac{3L}{4}$	$x=L$
$\psi=\sqrt{\frac{2}{L}}$	$\psi=0$



for $n=n \rightarrow n$ antinodes
 $\rightarrow n+1$ nodes

Ques:- What is the wavelength of de-Broglie wave attached with an e^- having speed (a) 300 m/sec (b) 3×10^7 m/sec.

Soln. we know $m_0 = 9.1 \times 10^{-31}$ kg.

(i) $V = 300$ m/sec. $V \ll c$

$$\therefore \lambda = \frac{h}{mV} = \frac{h}{\frac{m_0 V}{\sqrt{1 - \frac{V^2}{c^2}}}} \approx \frac{h}{m_0 V} \quad \left(\because \frac{V^2}{c^2} \approx 0 \right)$$

as $V \ll c$

$$= \frac{6.63 \times 10^{-34} \text{ J s}}{9.1 \times 10^{-31} \text{ kg} \times 300 \text{ m/sec.}} = 2.42 \times 10^{-3} \times 10^{-34} \times 10^{+31}$$

$$= 2.42 \times 10^{-6} \text{ m}$$

(ii) $V = 3 \times 10^7$ m/sec. \rightarrow comparable with c

$\therefore m \neq m_0$

$$m = \frac{m_0}{\sqrt{1 - \frac{9 \times 10^{14}}{9 \times 10^{16}}}}$$

$$m = \frac{m_0}{\sqrt{1 - 0.01}} = \frac{9.1 \times 10^{-31}}{0.99} = 9.2 \times 10^{-31} \text{ kg.}$$

$$\lambda = \frac{h}{mV} = \frac{6.63 \times 10^{-34}}{9.2 \times 10^{-31} \times 3 \times 10^7} = 0.24 \times 10^{-34 + 31 - 7} \text{ m}$$

$$\lambda = 0.24 \times 10^{-10} \text{ m}$$

To find ~~velocity~~ ^{wavelength} of electron with 80% speed of light.

Soln: Given $v = 80\% c$

$$= \frac{80}{100} \times 3 \times 10^8 = 0.8 c$$

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{m_0}{\sqrt{1 - 0.64}} = \frac{m_0}{\sqrt{0.36}}$$

$$m = \frac{m_0}{0.6} = \frac{9.1 \times 10^{-31}}{0.6} = 15.17 \times 10^{-31} \text{ kg.}$$

$$\lambda = \frac{h}{m v} = \frac{6.63 \times 10^{-34}}{15.17 \times 10^{-31} \times 0.8 \times 3 \times 10^8}$$

$$\lambda = 0.182 \times 10^{-11}$$